Learning from Tau Neutrino Appearance at Long Baselines

NuTau2021 Workshop, 29th September 2021 Kevin Kelly

$$\frac{1}{\sqrt{2}}\left(|\text{FNAL}\rangle + |\text{CERN}\rangle\right)$$

[1904.07265] (with de Gouvêa, Pasquini, Stenico) [2008.01088] (with Ellis & Li)

Outline

- Tau-neutrino Appearance at Long Baselines
- What can we learn from it?
 - "Standard" three-neutrino mixing
 - Additional physics beyond three neutrinos

Long-Baseline Oscillations

Long-Baseline, Three-Neutrino Oscillations

Assuming we're in the regime where the atmospheric mass-splitting dominates the solar one, the muon-neutrino to tau-neutrino oscillation probability can be *very* well approximated by

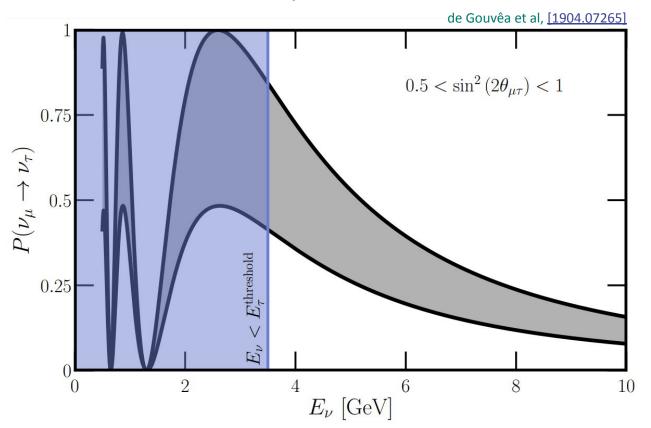
$$P(\nu_{\mu} \to \nu_{\tau}) = 4|U_{\mu 3}|^2|U_{\tau 3}|^2\sin^2\left(\frac{\Delta m_{31}^2 L}{4E_{\nu}}\right) + \text{subleading}$$

"Mixing angle" associated with tau-neutrino appearance, predicted to be large in the three-neutrino mixing paradigm

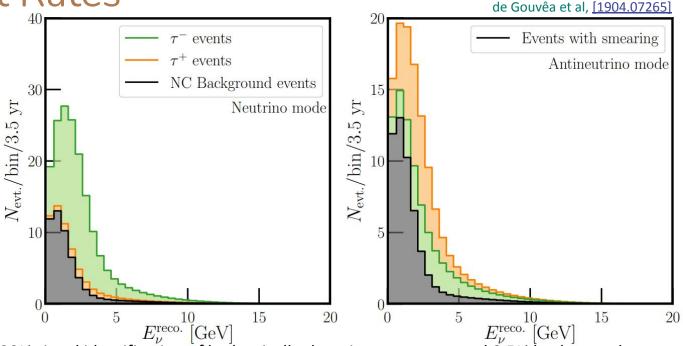
$$\sin^2 2\theta_{\mu\tau} \equiv 4|U_{\mu 3}|^2|U_{\tau 3}|^2$$

At the same baselines where muon-neutrino to electron-neutrino appearance is being studied (relatively small probability), the tau-neutrino appearance probability is large!

Appearance Probability at DUNE



Event Rates



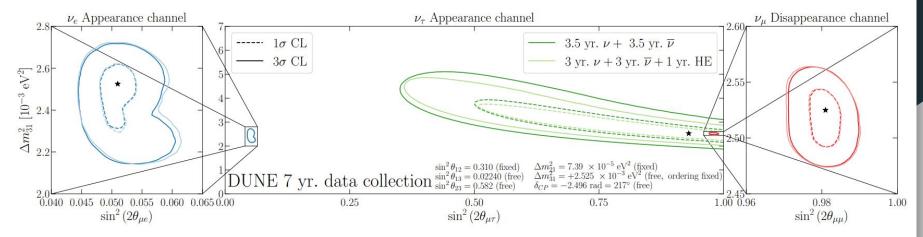
Considers 30% signal identification of hadronically-decaying tau events and 0.5% background contamination of neutral current events.

We also considered the "high-energy" beam tune -- results in higher event rate, but still just measuring the tail of the oscillation probability.

Extracting Three-Neutrino Physics

Three-Neutrino Measurements

de Gouvêa et al, [1904.07265]

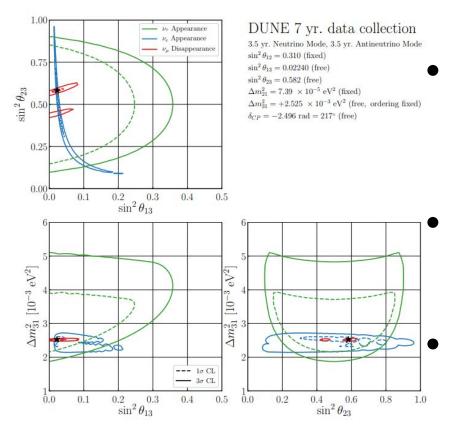


Within the three-neutrino picture, effective mixing angles are related, but we can determine the measurement capability of each of the three channels at DUNE. Consistency check -- do the mixing angles sum properly?

$$\sin^2 2\theta_{\mu e} \equiv 4|U_{\mu 3}|^2|U_{e 3}|^2, \qquad \sin^2 2\theta_{\mu \mu} \equiv 4|U_{\mu 3}|^2(1-|U_{\mu 3}|^2)$$

$$\sin^2 2\theta_{\mu e} + \sin^2 2\theta_{\mu \tau} = \sin^2 2\theta_{\mu \mu}$$

Independent Mixing/Mass Splitting Measurements

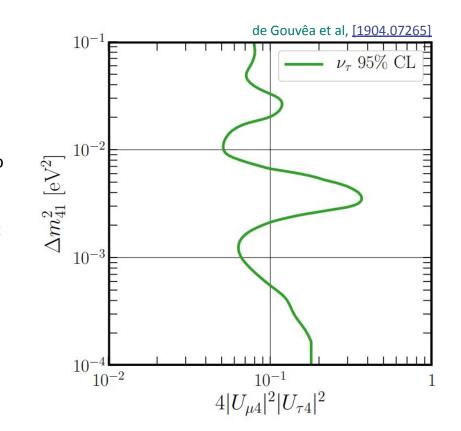


Separating out oscillation channels in the analysis, attempting to independently measure the mixing angles (and CP violation, not shown), as well as the atmospheric mass-squared splitting. Unsurprisingly, the tau channel (green) is weaker than the combined muon disappearance (red) and electron appearance (blue) channels. Still, this can serve as a consistency check.

Some Physics Beyond Three Neutrinos

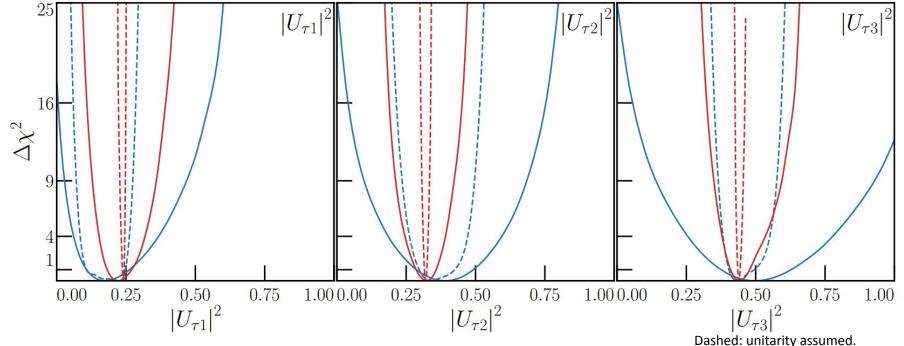
Light Sterile Neutrino

In the 3+1 Sterile Neutrino scenario, tau neutrino appearance has sensitivity to one of the three new mixing angles that electron neutrino appearance/muon neutrino disappearance don't (except through matter effects).









Unsurprisingly, the tau row of the Leptonic Mixing Matrix is the least understood. If we abandon the assumption of unitarity, we need direct measurements of tau neutrinos -- DUNE, and similarly, IceCube Gen2, have a lot to offer with tau neutrino appearance.

Dasned: unitarity assumed.

Blue: Current data

Red: Future (+DUNE, IceCube Gen2)

Conclusions

- With upcoming experiments, the future of tau-neutrino appearance is bright!
- While this information serves more as a "consistency check" on three-neutrino physics, it still has much to offer in well-motivated extensions beyond three neutrinos.